

SPACEWATCH® Observations of Asteroids and Comets Supporting the Large- Scale Surveys.

Poster 414.11

46th Mtg of the DPS/AAS, 2014 Nov.

Robert S. McMillan¹, T. H. Bressi¹, J. V.
Scotti¹, J. A. Larsen², R. A. Mastaler¹, and
A. F. Tubbiolo¹

URL: <http://spacewatch.lpl.arizona.edu>

¹University of Arizona; ²U.S. Naval Academy

Photo by Marcus L. Perry, 1997

Summary

- Follow-up of "large" NEOs ($H \leq 22$) as they recede from Earth after discovery and become fainter, as well as VIs, PHAs, & NEOs observed by WISE.
- New, faster-reading CCD on 1.8-meter telescope.
- Observed at elongations as small as 46° .
- ~2800 tracklets of NEOs from Spacewatch accepted by MPC each year.
- Big, long archive from 0.9-m telescope to support precoveries.

Why Targeted Followup is Needed

- Discovery arcs too short to define orbits:
 - Followup observation intervals need to be thousands of times longer than discoveries.
- Objects can escape redetection by surveys:
 - Surveys too busy covering other sky.
 - Objects tend to get fainter after discovery.
- Sky density of detectable NEOs is too sparse to rely on incidental redetections alone.

Why More Followup is Needed

- 1/3rd of PHAs observed on only 1 opposition.
- 1/6th of PHAs' arcs $< 30^{\text{d}}$.
- ~Half of potential close approaches in the next 30 years will be by objects observed on only one opposition.
- 2/3^{rds} of $H \leq 22$ VI's on JPL risk page *are lost* and $>$ half of those were discovered within the last 6 years by modern surveys.

How “lost” can they get?

- (719) Albert discovered visually in 1911.
- “Big” Amor asteroid, diameter ~ 2 km.
- Favorable (perihelic) apparitions 30 yrs apart.
- Missed in 1941 due to inattention.
- Missed in 1971 due to large uncertainty.
- MPC recognized (719) as a rediscovery by Spacewatch in 2000.

1979 XB: A “Big” Lost “VI”!

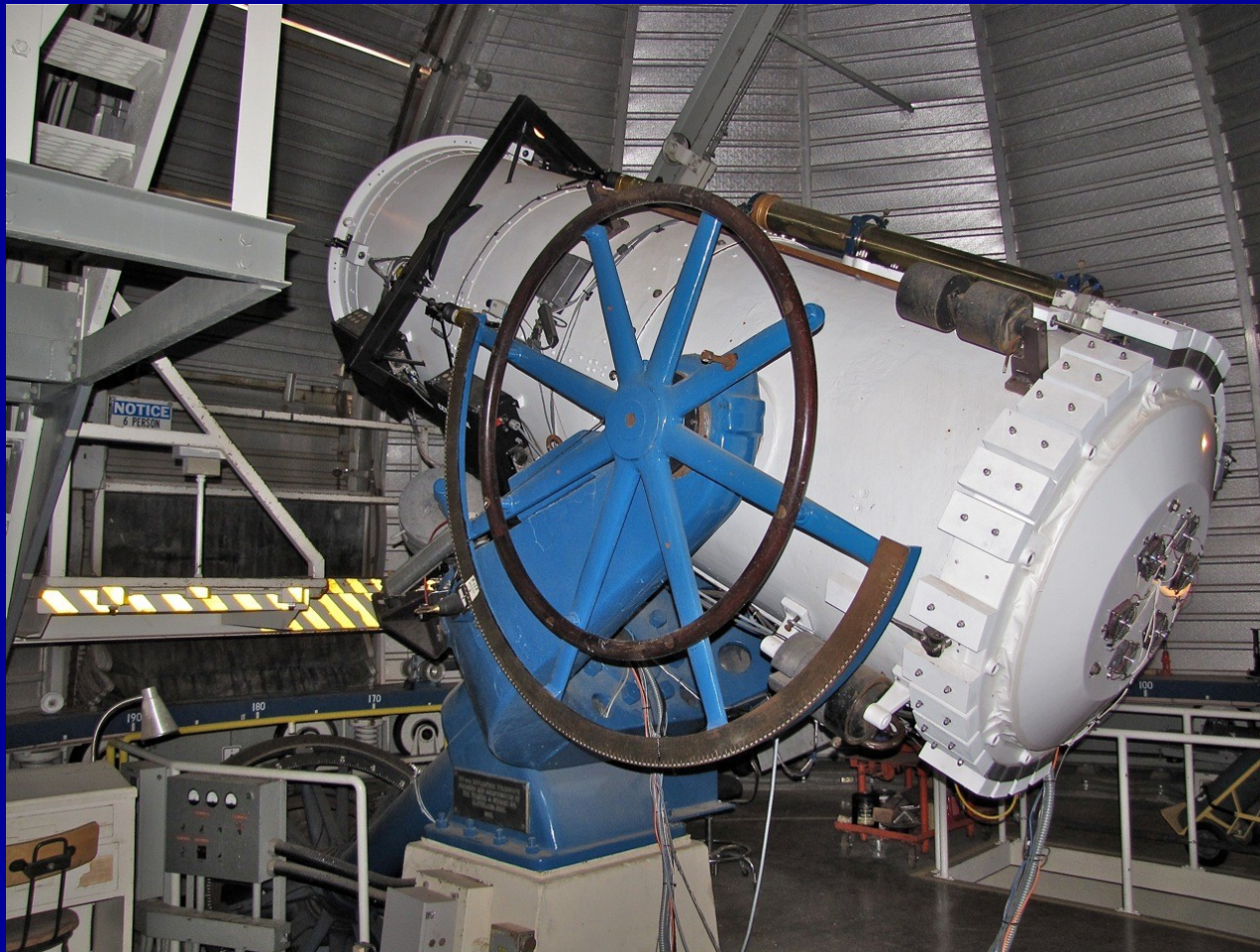
- 4-day observed arc in 1979 December.
- $H \approx 18.5 \leftrightarrow$ Diameter 370-1200 m.
- Synodic period $\approx 1.4_y$.
- Possible close encounters in 2056 & 2086.
- Not rediscovered in >3 decades of modern surveying.

0.9-m Telescope Modernized by Spacewatch in 2002

- Hyperboloidal primary & refractive field corrector.
- Mosaic of 4 CCDs.
- Bandpass $\approx 0.5\text{-}0.9\ \mu\text{m}$; $\lambda_{\text{eff}} \approx 0.7\ \mu\text{m}$.
- Began 2003 April; 24 nights per lunation.
- Automated in 2005 May.
- Patterns near opp'n, & low elongation in east.
- $1400\ \text{deg}^2$ per lunation; $V\ \text{mag} \approx 20.5\text{-}21.7$.

0.9-m Telescope; Site Code 691

Photo by Roger Carpenter, 2012 Feb



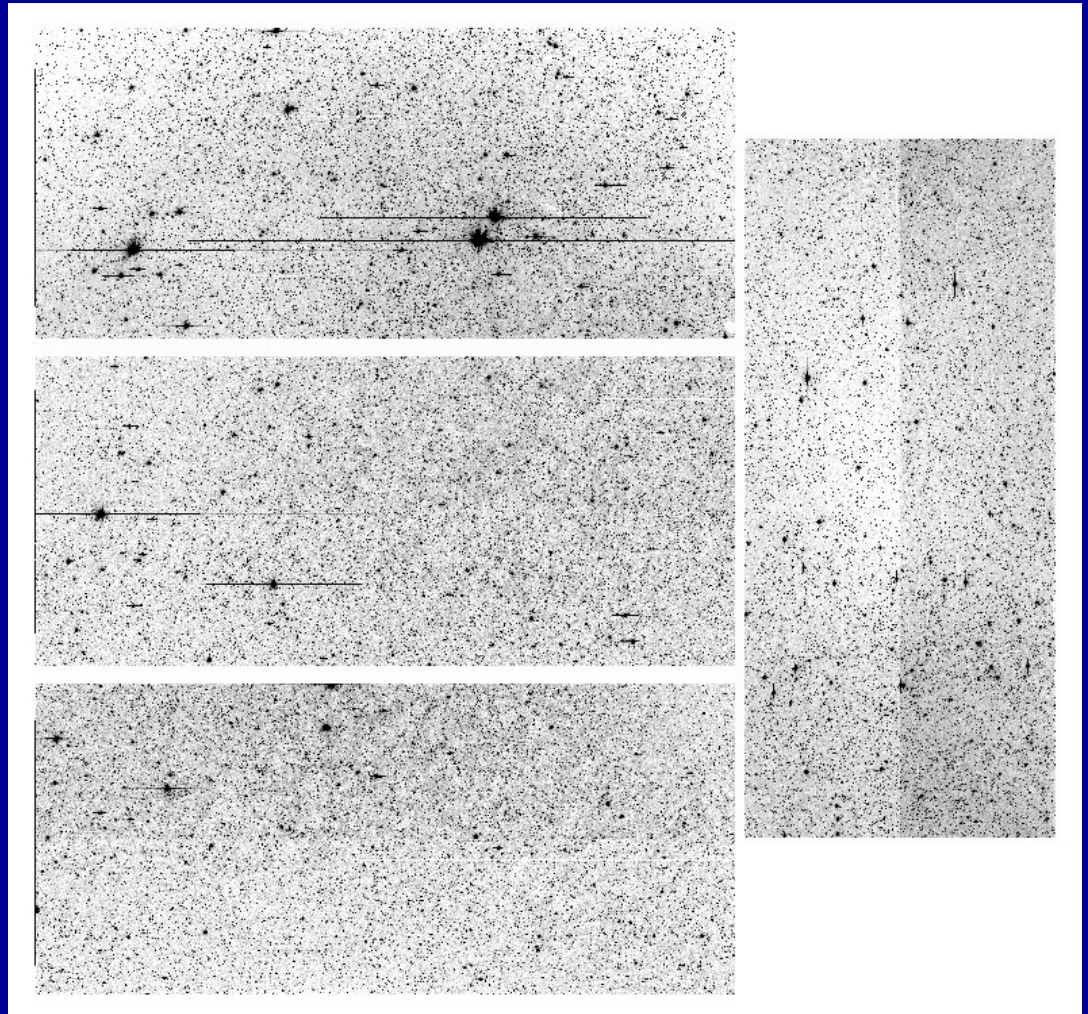
**Spacewatch CCD Mosaic
on 0.9-m telescope.**

Four EEV Grade-1,
back-illuminated,
antireflection-coated
CCDs of 4608x2048
pixels each.

37 million pixels.

1 arcsec per pixel.

2.9 deg² covered.



Archive from Mosaic on 0.9-m:

- Revisits @ 4^d intervals aid MBA linkages.
- ~20 TB in size.
- 11.5 yrs of uniformly conducted surveying.
- Incidental astrometry & precoveries of NEOs.
- V mag limit ~20-21.
- Coverage ~1400 deg² per lunation (3 passes per pointing) mostly along ecliptic and low-elongation in the east.

**Spacewatch 1.8-meter
Telescope on Kitt Peak;
Site code 291**

New CCD in 2011 Oct:

FOV = $20' \times 20'$.

Scale = 0.6 arcsec/pixel.

Bandpass \square “V+R+I”.

Faster readout.

Limit V=23.

50% more obs per year.

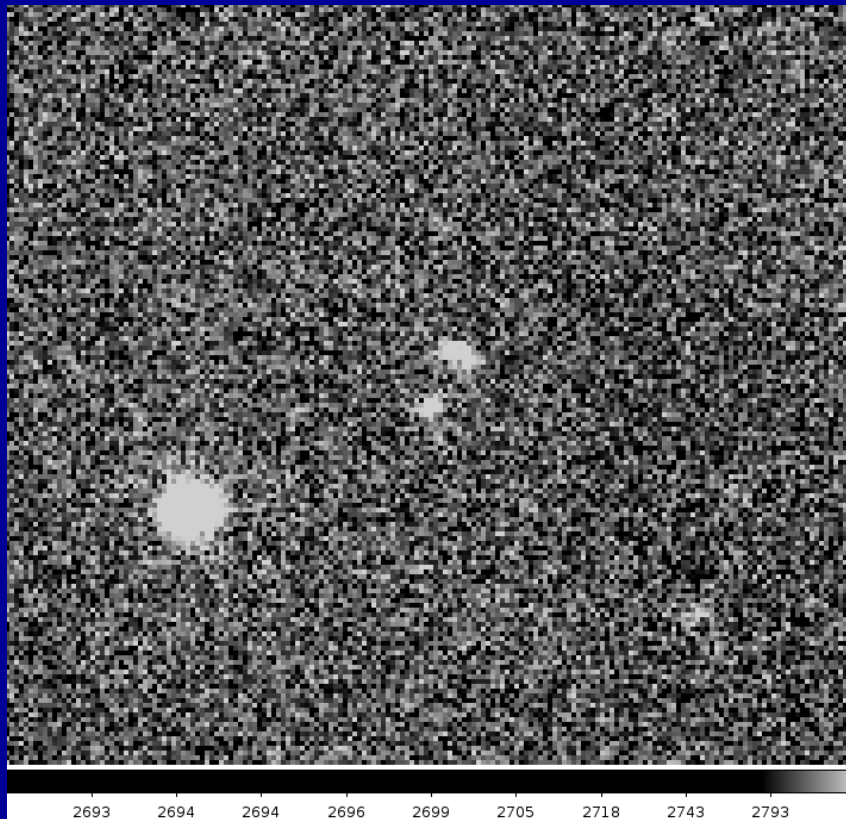
Astrometric resids \square **0.3
arcsec**, vs. 0.6 on NEOs
with the old CCD.

Photo by Roger Carpenter, 2012 Feb.

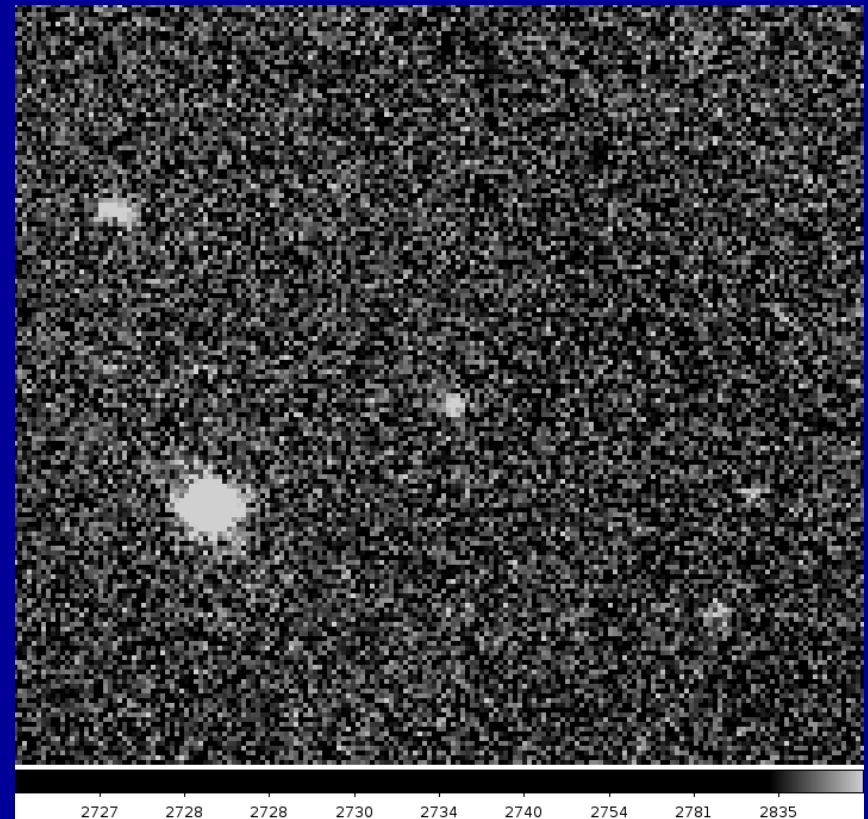


2014 SE145 on 2014 Sep 23, Spacewatch 1.8-m.
R=18, 5 sec exp.

12:11:29 UT



12:12:21 UT



2014 SE 145: MPEC 2014-S75

Orbital elements: Earth MOID = 0.01 AU

a=2.2 AU e=0.55 Incl=8 deg P=3.33 yr

AbsMag = 27.5V Orbital element Uncertainty U=6

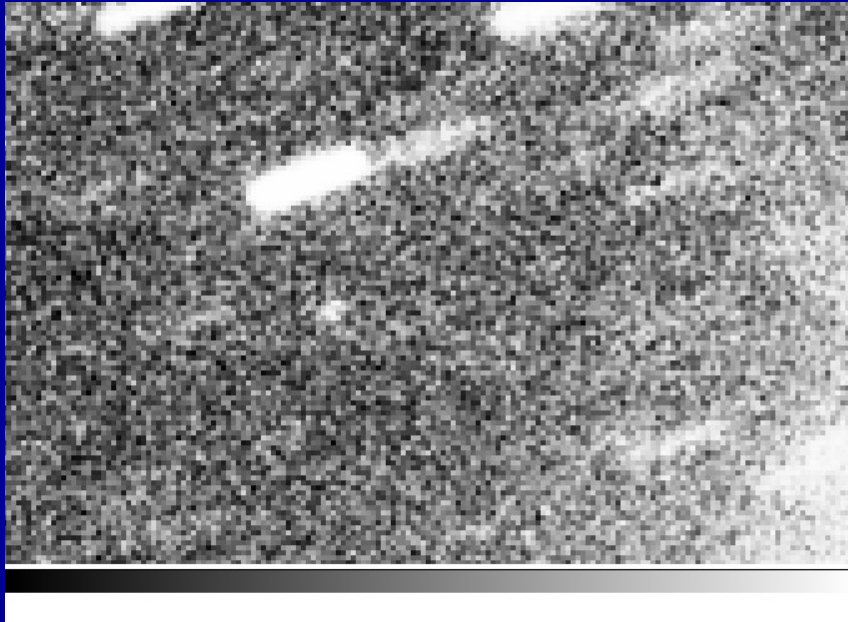
Residuals in seconds of arc

140923	I41(-12.4 +6.5)	140923	291	-0.2	+0.1	140924	474	-0.4	+0.3
140923	I41(-16.2 +6.4)	140924	104	0.0	-0.2	140924	474	-0.3	-0.2
140923	291 +0.1 -0.1	140924	104	+0.1	+0.1	140924	474	+0.5	-0.2
140923	291 0.0 0.0	140924	104	-0.2	+0.1	140924	474	+0.2	+0.1

2014 SN1 (F51 disc., formerly a VI),
diam=20-80 m, $V=22.5$, 0.7 arcsec/min.

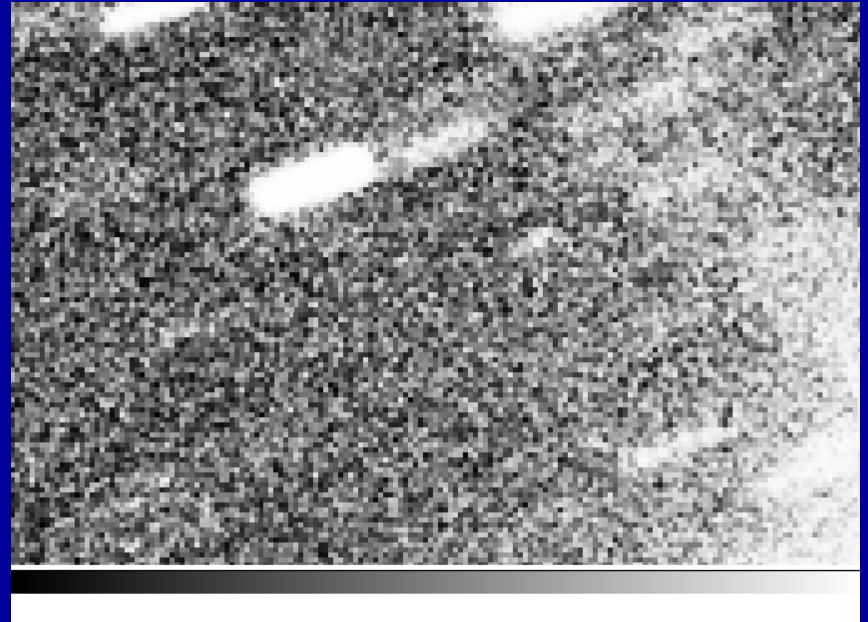
2014 Oct 1 04:29:04

14x120 sec stacked exp's.



2014 Oct 1 05:06:32

14x120 sec stacked exp's.



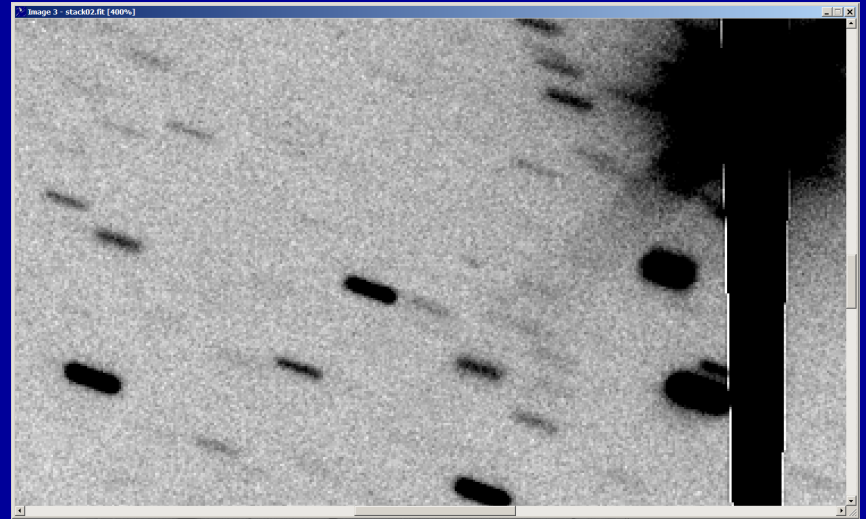
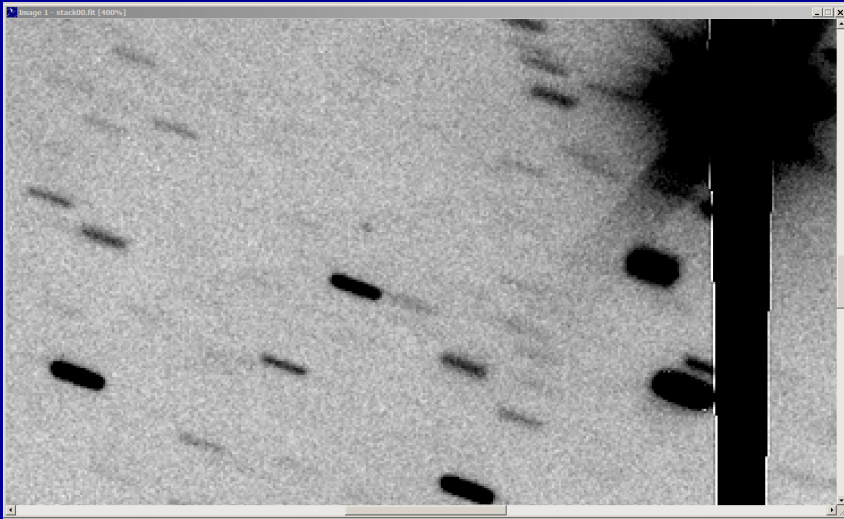
2014 SR261 (G96 discovery, VI),
diam=7-30 m, $V=22.5$, 0.2 arcsec/min.

2014 Oct 1 07:55:00

14x120 sec stacked exp's.

2014 Oct 1 08:32:28

14x120 sec stacked exp's.



2014 RS17: Former VI, disc COD 703, $R=22.5$,
diam=40-150 m, 0.6 arcsec/min, 2014 Oct 1 06:00:34.
18 stacked exp's @120 sec.



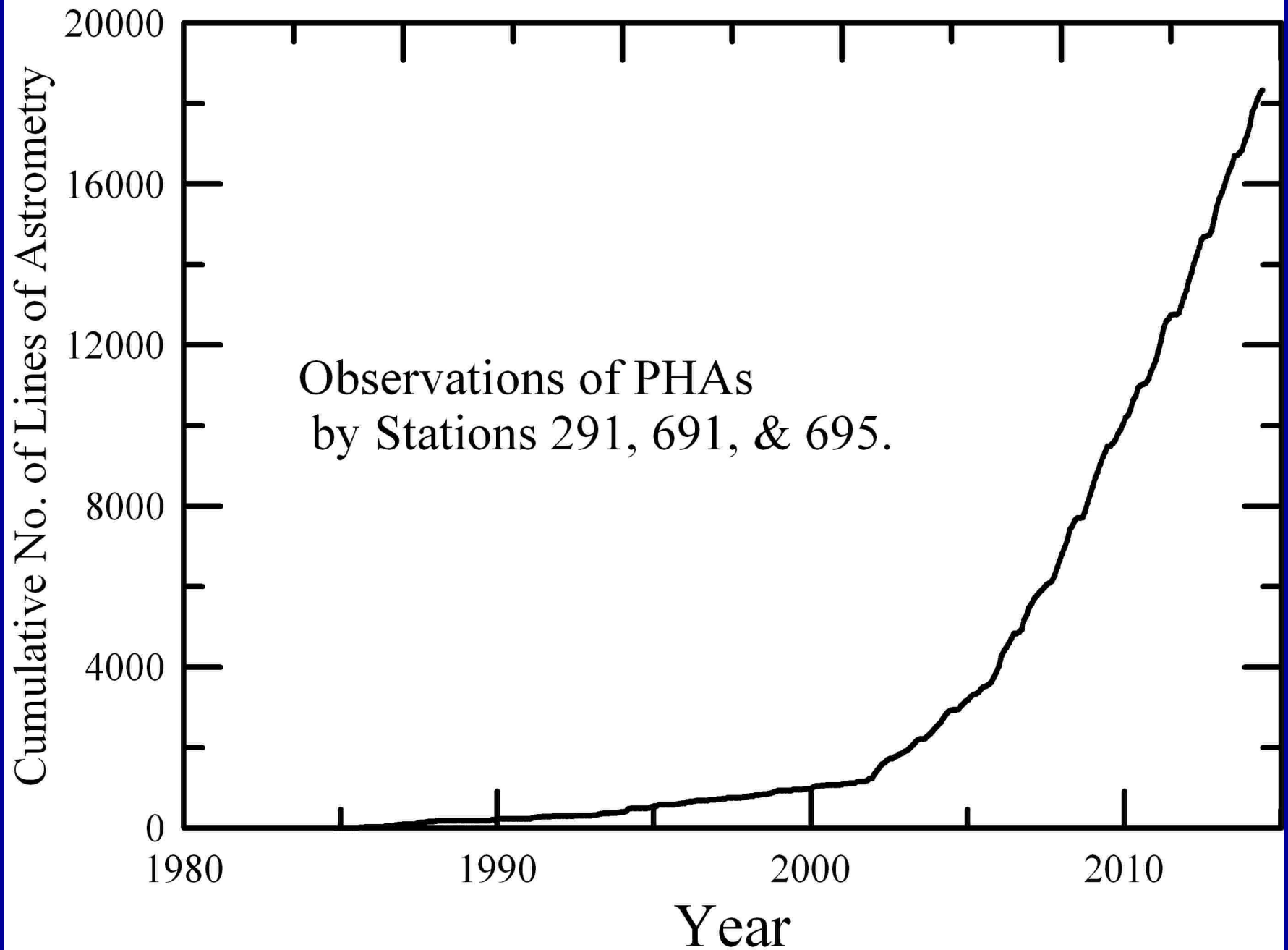
Recent Spacewatch Results

- Annual average of 8,492 lines of astrometry of 1,018 different NEOs including 177 different PHAs /yr.
- Contributed to the removal of half of the objects that were retired from JPL's impact risk list.
- Per year we observe...
 - about 35 radar targets,
 - 50 NEOs that were measured by NEOWISE, and
 - 100 potential rendezvous destinations.

Spacewatch Contributions

- Between 2011 Oct 16 & 2014 Sep 25
Spacewatch observed:
 - 50% of all NEOs observed in that time.
 - 54% of all PHAs observed in that time.
- Leading station in followup of provisionally designated PHAs while faint ($V \geq 22$);
contributing 41% of all such observations.

Cumulative Spacewatch Observations of PHAs



Needs of Followup Campaign (besides money).

- Get longer arcs during discovery apparitions:
 - Require longer arcs for posting on NEO CP.
 - Keep provisional desigs. on the NEO CP.
 - Help us get more time on larger telescopes.
- More focused selections of targets in MPC's NEA Observation Planner:
 - NEOWISE, NHATS, Radar, comet candidates.

Acknowledgements

- NASA's NEO Observation Program.
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- JPL's NEO Office.
- The U. S. Naval Academy.
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- Other private donors.